Release 3

Data for neutrino fluxes at production

The data for neutrino fluxes from DM annihilations in the Earth and in the Sun, computed in

M. Cirelli, N. Fornengo, T. Montaruli, I. Sokalski, A. Strumia, F. Vissani "Spectra of neutrinos from dark matter annihilation" hep-ph/0506298v4

can be downloaded from any of the following locations:

- www.to.infn.it/~fornengo/DMnu.html
- www.cern.ch/astrumia/DMnu.html
- www.marcocirelli.net/DMnu.html

The neutrino fluxes at production are given in two forms:

- 1. as raw numerical tables of the fluxes;
- 2. as tables of fit parameters.

1. Numerical tables of the fluxes

The file DMnuProdFluxes.zip contains the following data files:

earth_numu.dat	fluxes of $\nu_{\mu} = \bar{\nu}_{\mu} = \nu_e = \bar{\nu}_e$ from annihilations in Earth
earth_nutau.dat	fluxes of $\nu_{\tau} = \bar{\nu}_{\tau}$ from annihilations in Earth
sun_numu.dat	fluxes of $\nu_{\mu} = \bar{\nu}_{\mu} = \nu_e = \bar{\nu}_e$ from annihilations in Sun
sun_nutau.dat	fluxes of $\nu_{\tau} = \bar{\nu}_{\tau}$ from annihilations in Sun

Each file is a table of the following format:

 $m_{\rm DM}$ x b τ c q t Z W Z(no prompt) W(no prompt)

where $m_{\rm DM}$ is the mass of the annihilating DM particle, $x = E_{\nu}/m_{\rm DM}$ is the rescaled neutrino energy and the other columns give the neutrino fluxes normalized per single DM annihilation for the annihilation channels $b\bar{b}$, $\tau\bar{\tau}$, $c\bar{c}$, light quarks ($u\bar{u} \equiv d\bar{d} \cong s\bar{s}$), ZZ, W^+W^- . The last two column give the neutrinos fluxes from the ZZ and W^+W^- channels having subtracted the "prompt" neutrinos (see the text of the paper). The values of $m_{\rm DM}$ provided are the following. For the case of the Earth (14 values):

 $m_{\rm DM} = \{10, 30, 50, 70, 90, 100, 150, 200, 250, 300, 500, 700, 900, 1000\}$ GeV.

For the case of the Sun (12 values):

 $m_{\rm DM} = \{10, 30, 50, 70, 90, 100, 200, 300, 500, 700, 900, 1000\}$ GeV.

2. Tables of fit parameters

The file DMnuProdParameters.zip contains the following data files:

parameters_fit_numu_earth.dat

fit parameters for $\nu_\mu=\bar\nu_\mu=\nu_e=\bar\nu_e$ from annihilations in the Earth parameters_fit_nutau_earth.dat

fit parameters for $\nu_\tau=\bar\nu_\tau$ from annihilations in the Earth <code>parameters_fit_numu_sun.dat</code>

fit parameters for $\nu_{\mu} = \bar{\nu}_{\mu} = \nu_e = \bar{\nu}_e$ from annihilations in the Sun parameters_fit_nutau_sun.dat

fit parameters for $\nu_{\tau} = \bar{\nu}_{\tau}$ from annihilations in the Sun

Each file provides the fitting parameters a_i , b and c_i to be used in the expression for the neutrino flux

$$\frac{dN}{dx} = g(x) = a_0(1 + a_1w + a_2w^2 + a_3w^3 + a_4w^4 + a_5w^5)(1 - x)^b + c_0x^{c_1}(1 - x)^{c_2}$$
$$x = E_{\nu}/m_{\rm DM}, \qquad w = \log_{10}x,$$

which corresponds to eq. (7) in the paper. The parameters are given as a table in the format

 $m_{\rm DM}$ a_0 a_1 a_2 a_3 a_4 a_5 b c_0 c_1 c_2

(where $m_{\rm DM}$ is the mass of the annihilating DM particle) for each of the following annihilation channels

{ $b\bar{b}, \tau\bar{\tau}, c\bar{c}$, light quarks $(u\bar{u} \equiv d\bar{d} \cong s\bar{s})$, gluons }.

The values of $m_{\rm DM}$ provided are the following. For the case of the Earth (14 values):

 $m_{\rm DM} = \{10, 30, 50, 70, 90, 100, 150, 200, 250, 300, 500, 700, 900, 1000\}$ GeV.

For the case of the Sun (11 values):

 $m_{\rm DM} = \{10, 30, 50, 70, 90, 100, 300, 500, 700, 900, 1000\}$ GeV.

These data correspond to the data provided in Table 1 and Table 2 of the paper. See further discussion there.

Differences with Release 2

A numerical bug in the implementation of the boost for top quark decays has been fixed (modifications are quite small; they affect the $t\bar{t}$ channels in Figures 2, $5 \rightarrow 12$, as well as in Tables 3 and 4 in the paper).

The present Release 3 corresponds to v4 of the paper in the arXiv.

Differences between Release 2 and Release 1

An erroneous double counting of the prompt neutrino yield in W-boson decays has been fixed (modifications affect the W^+W^- and $t\bar{t}$ channels in Figures 2, $5 \rightarrow 12$, as well as in Tables 3 and 4 in the paper) and a few parameters have been updated (modifications are generally small or null).

Release 2 corresponds to v3 of the paper in the arXiv. Previous versions on the arXiv and the journal version on Nuclear Physics B had used Release 1.