

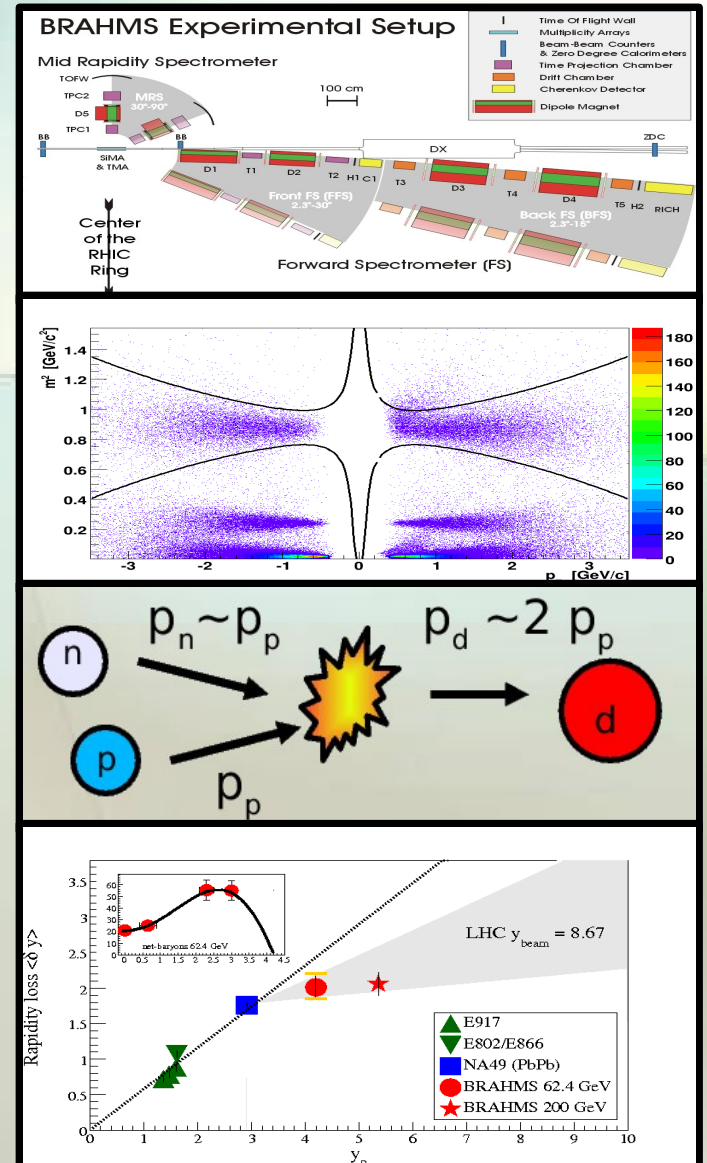


Recent Results from the BRAHMS Collaboration - Deuteron Coalescence and Nuclear Stopping

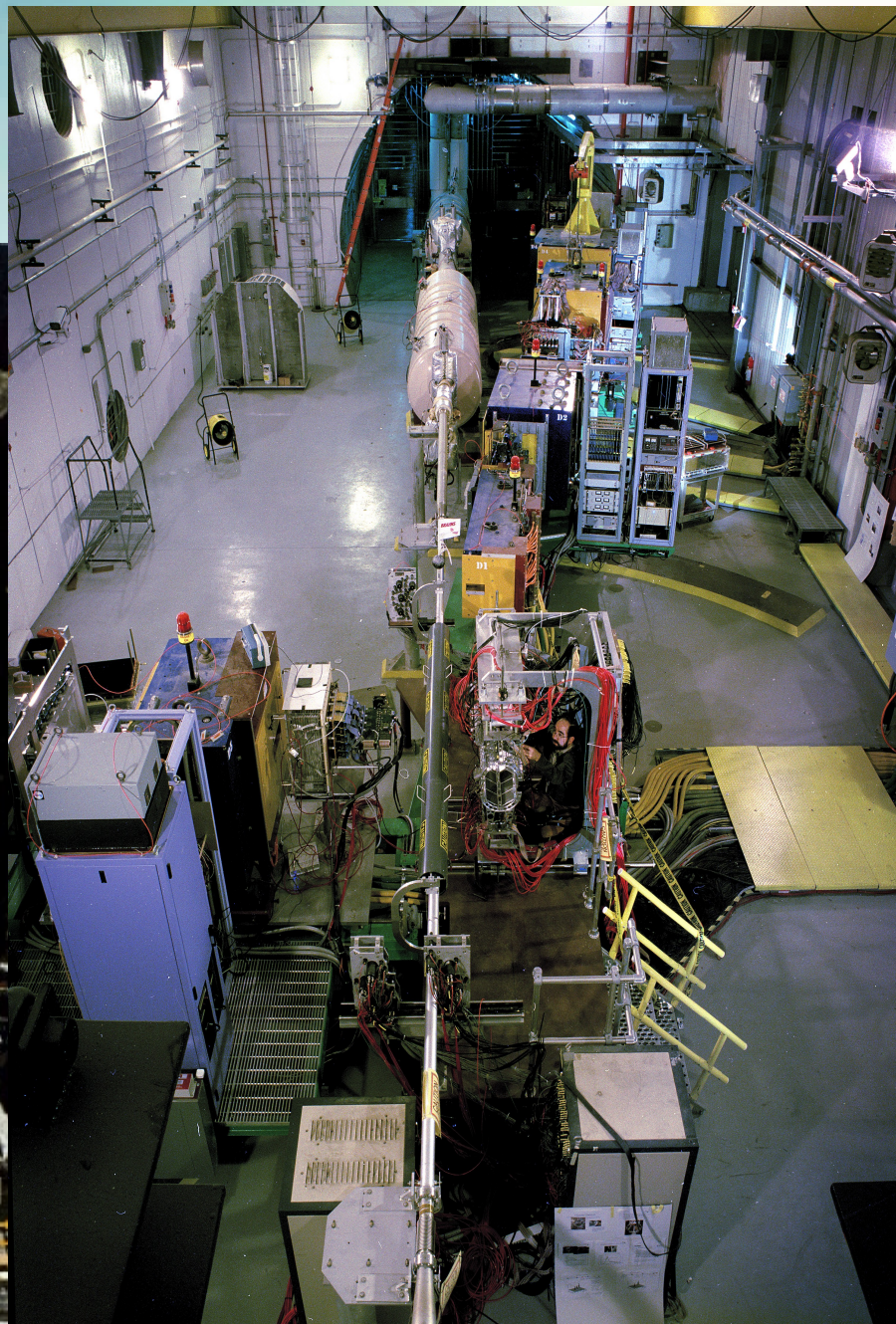
Casper Nygaard, Niels Bohr Institute
BRAHMS Collaboration

Outline

- The Brahms Experiment
- Analysis
- Coalescence Results.....
- Nuclear Stopping Results.....

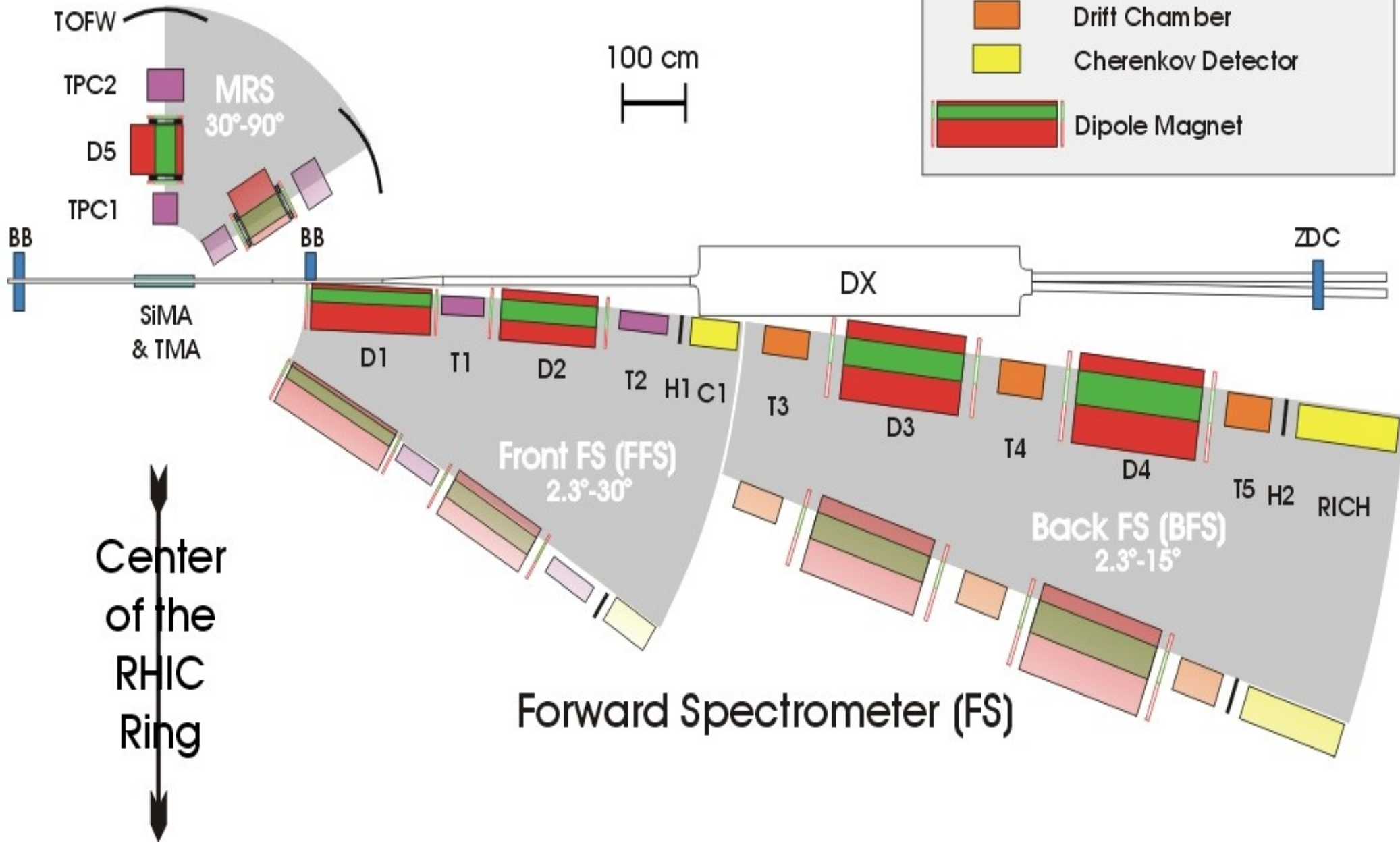


BRAHMS @ RHIC

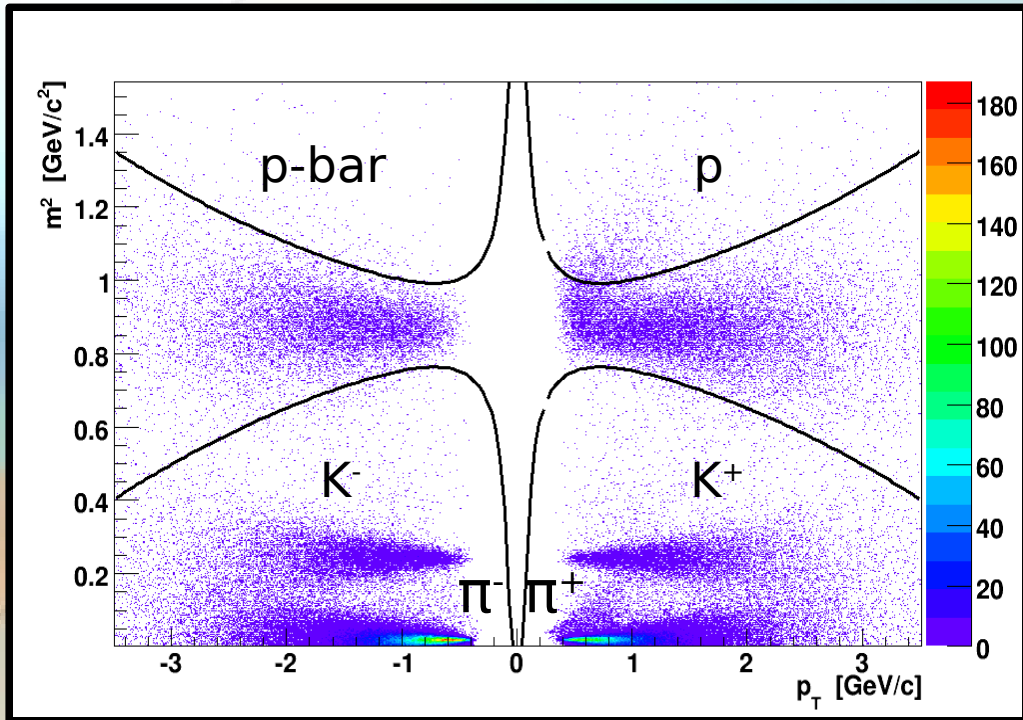


BRAHMS Experimental Setup

Mid Rapidity Spectrometer

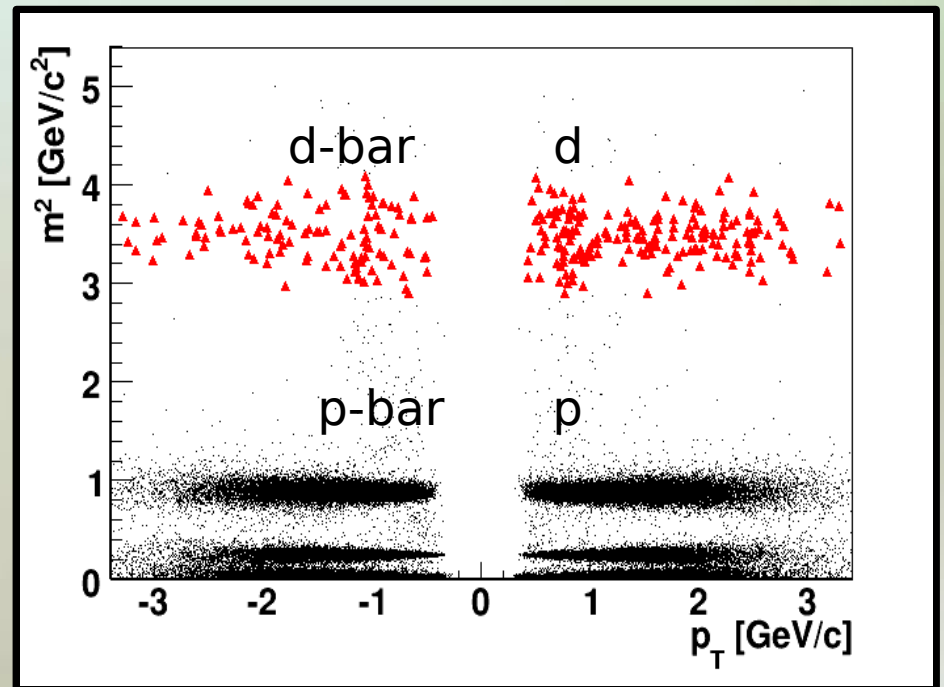


TOF PID



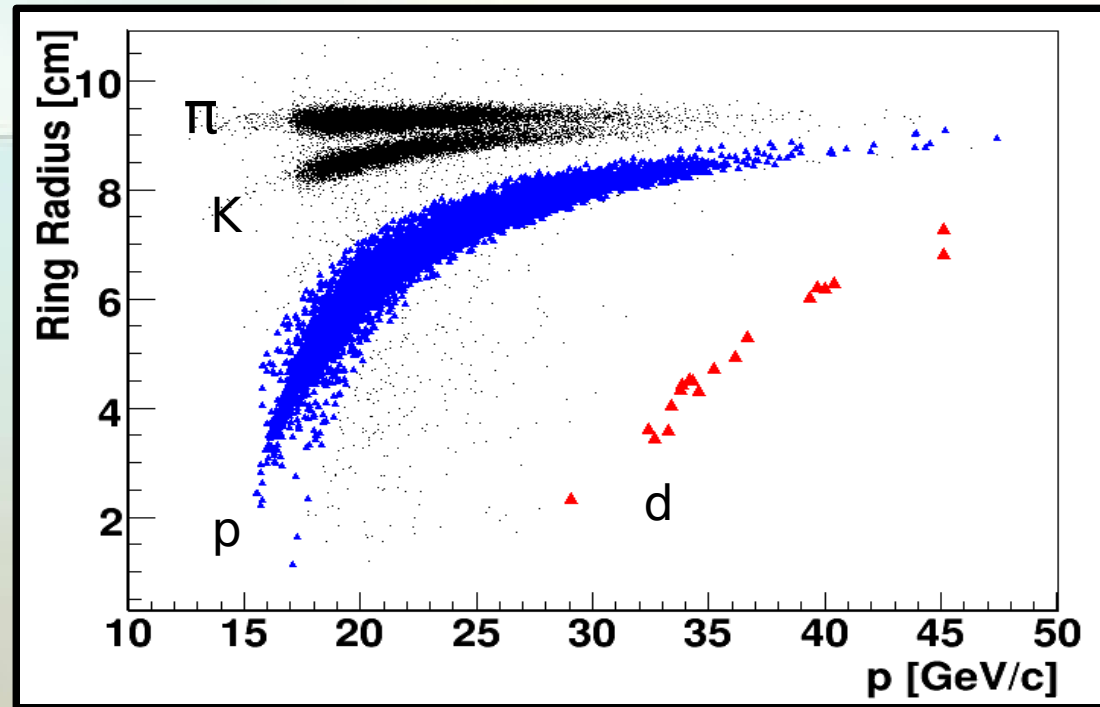
- TOF PID used in the MRS and at $y \sim 2$.

- Proton PID done by fitting the m^2 vs. p_T distribution.
- Deuteron PID done by a gaussian fit in the m^2 distribution.



RICH PID

- Proton PID
 - Direct: From $p \sim 15$ GeV/c, the Cherenkov ring radius is used
 - Indirect: $12 > p > 17$ GeV/c
- Deuteron PID
 - Direct: From $p \sim 30$ GeV/c
- Used for PID at $y \sim 3$.

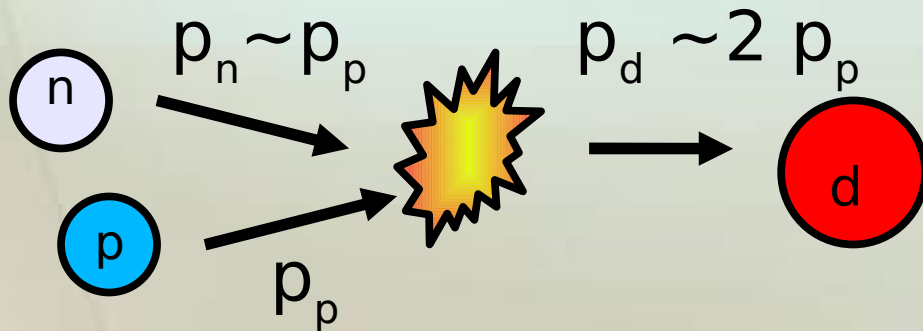


Spectra

- The invariant spectra have been corrected for:
 - Acceptance
 - Tracking efficiency
 - Multiple scattering, absorption and weak decay for (anti)-protons by GEANT
 - GEANT does not handle anti-deuterons, and does not handle hadronic interactions for deuterons.
 - Deuteron correction approximated to:
$$\text{Eff}(p_{d/d\bar{d}}) = \text{Eff}(p_d)_{\text{GEANT}(d)} * (\text{Eff}(p_d/2)_{\text{GEANT:hadronic}(p/p\bar{p})})^2$$

Coalescence

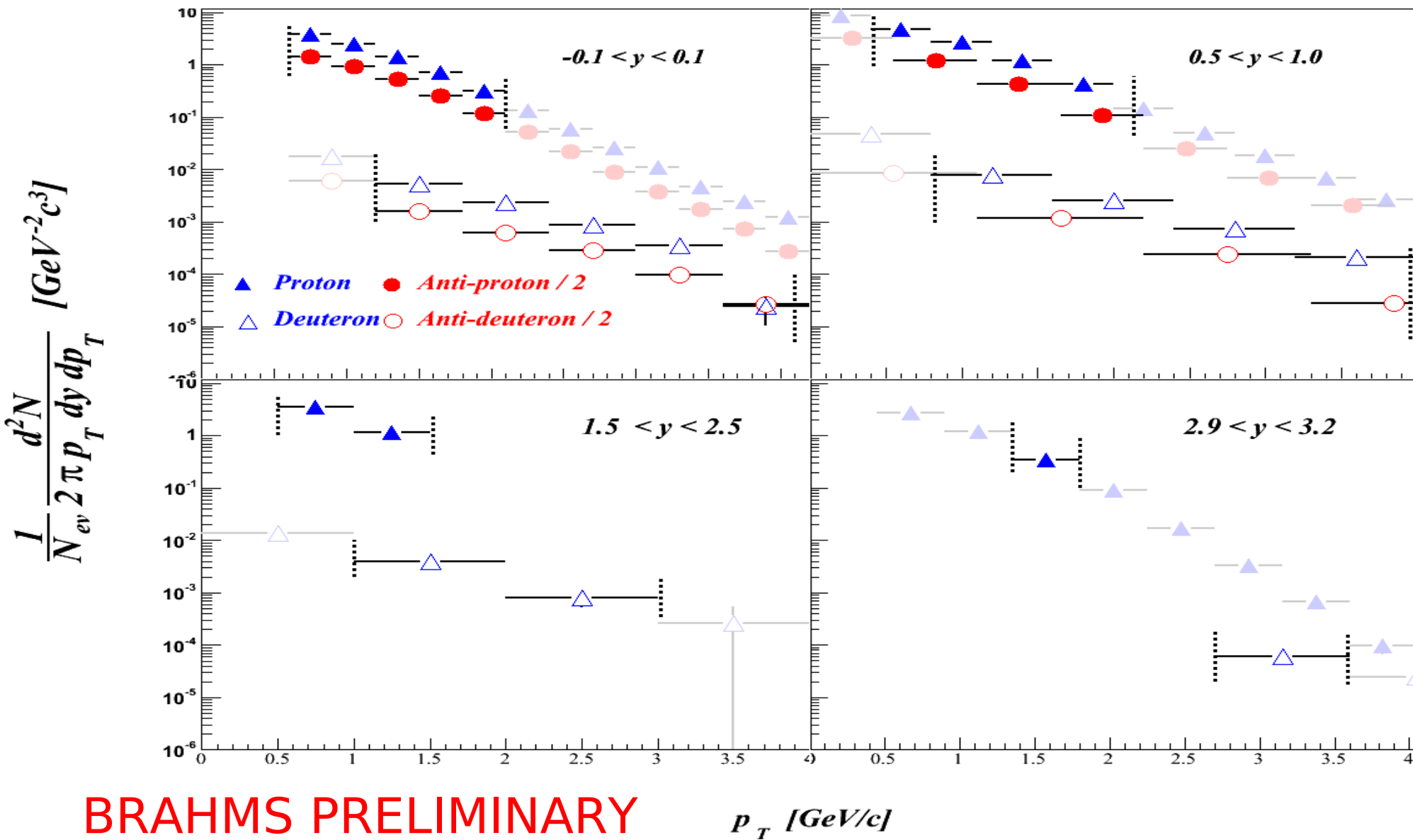
- Deuteron coalescence is the creation of a deuteron, from a proton and a neutron.
- Due to the very low binding energy of the deuteron (2.22 MeV), Coalescence probes the collision at the timescale of the freeze-out.
- Coalescence parameter given by:



$$B_2 = \frac{E_d \cdot \left(\frac{d^3 N_d}{dp_d^3} \right)}{\left(\frac{E_p \cdot d^3 N_p}{dp_p^3} \right)^2}$$

- B_2 is inversely proportional to the collision volume according to various models. [Pearson]

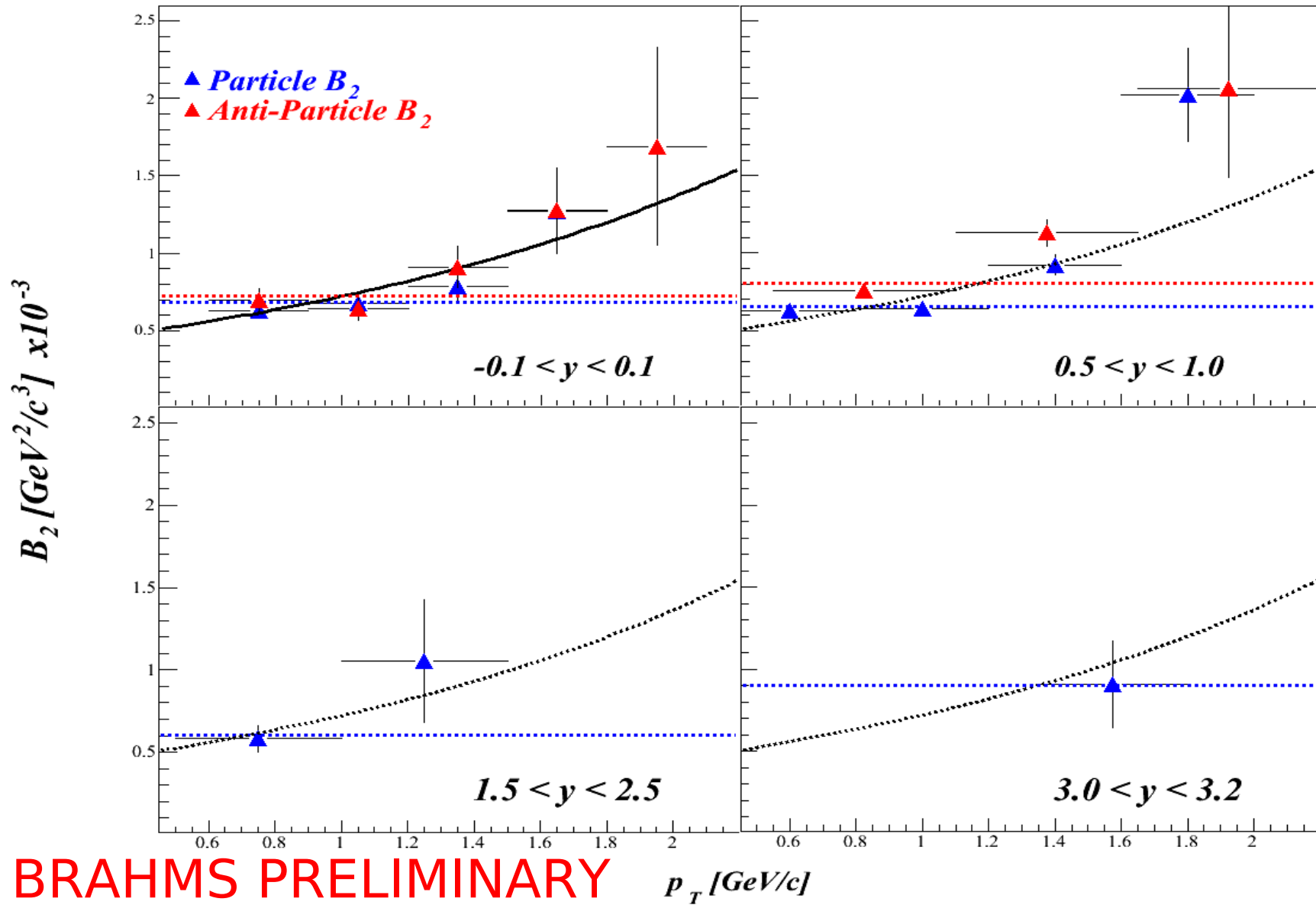
Spectra (0-20% central Au-Au@200GeV)



BRAHMS PRELIMINARY

p_T [GeV/c]

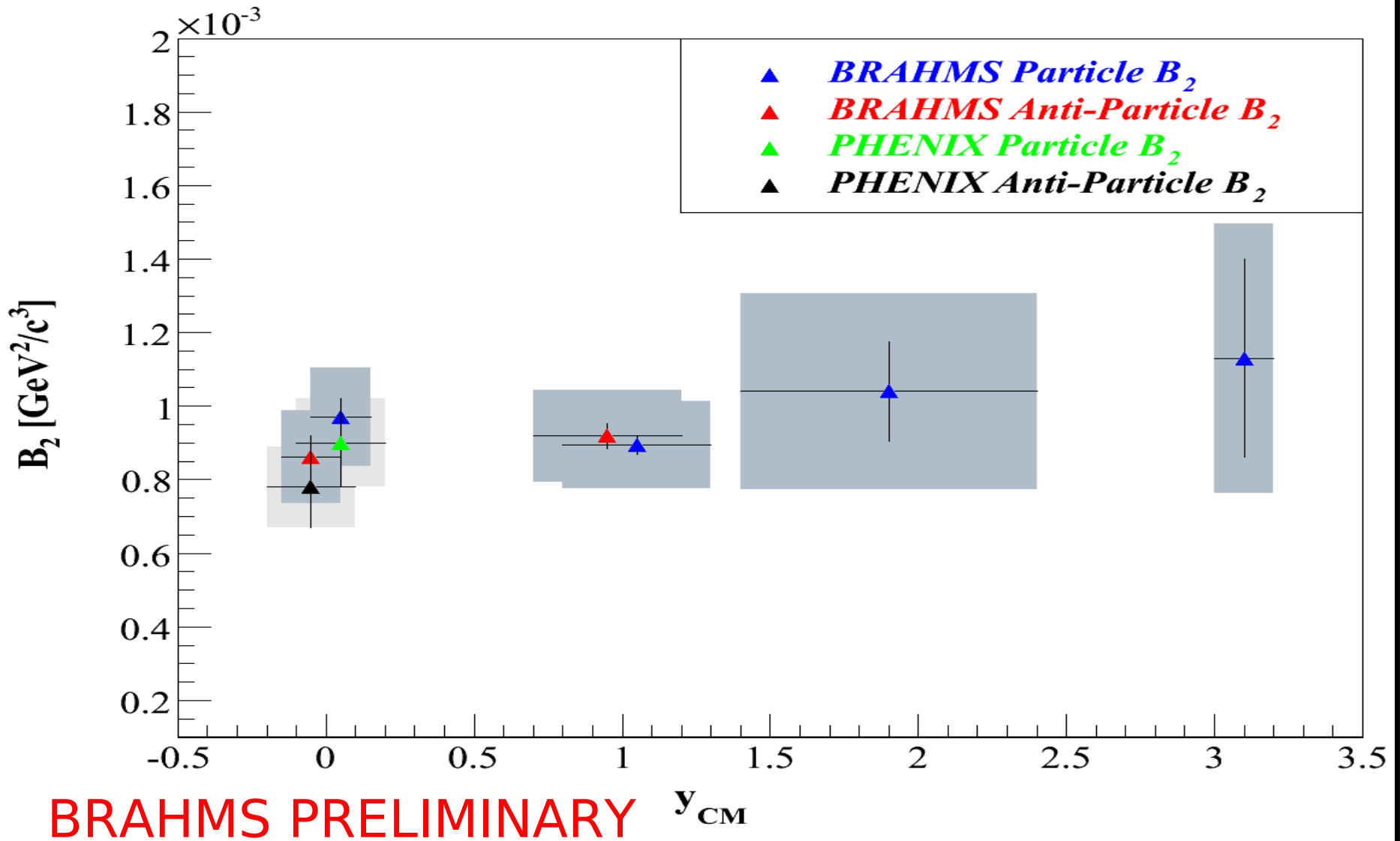
B_2 vs. p_T



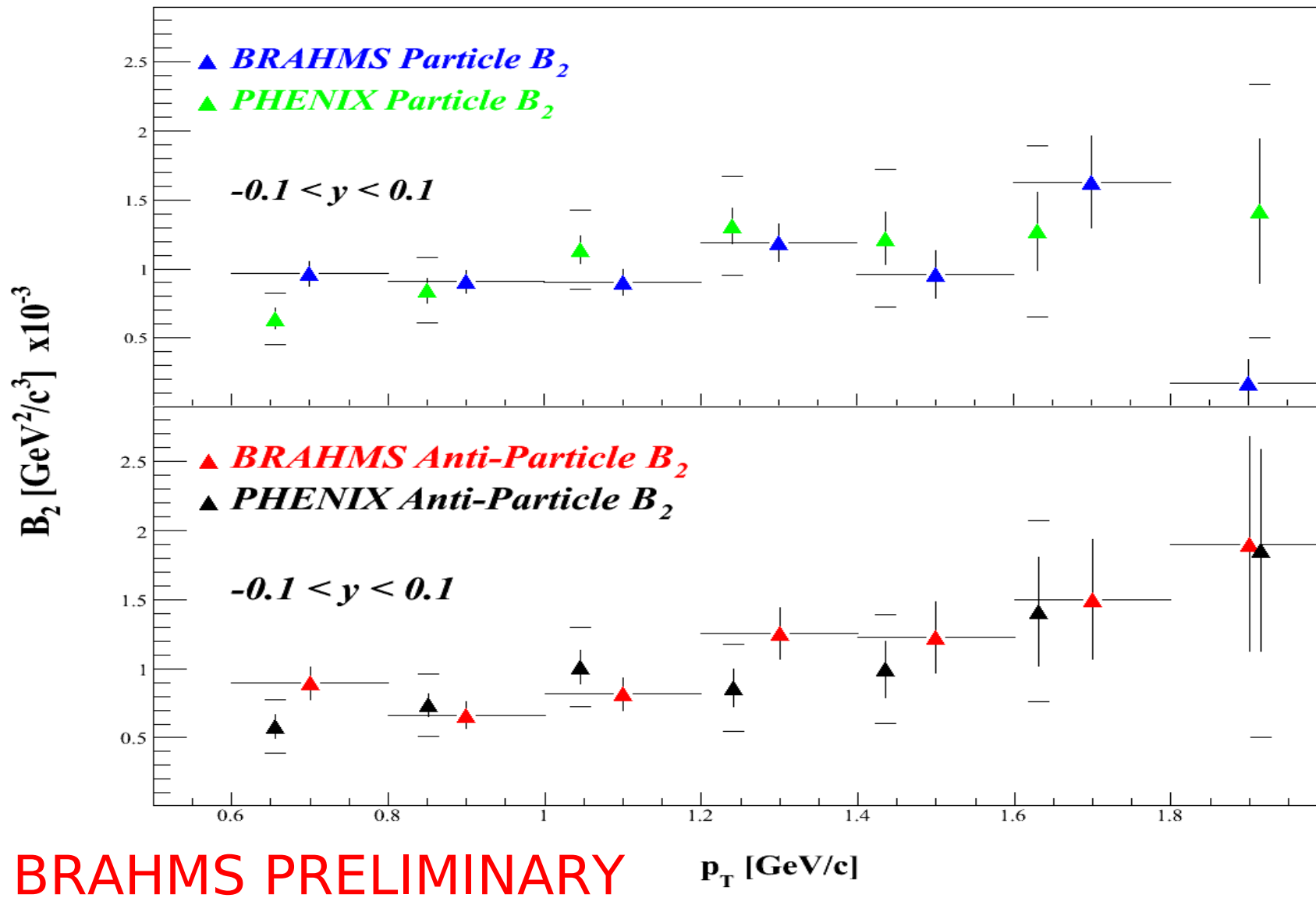
BRAHMS PRELIMINARY

p_T [GeV/c]

B_2 vs. y

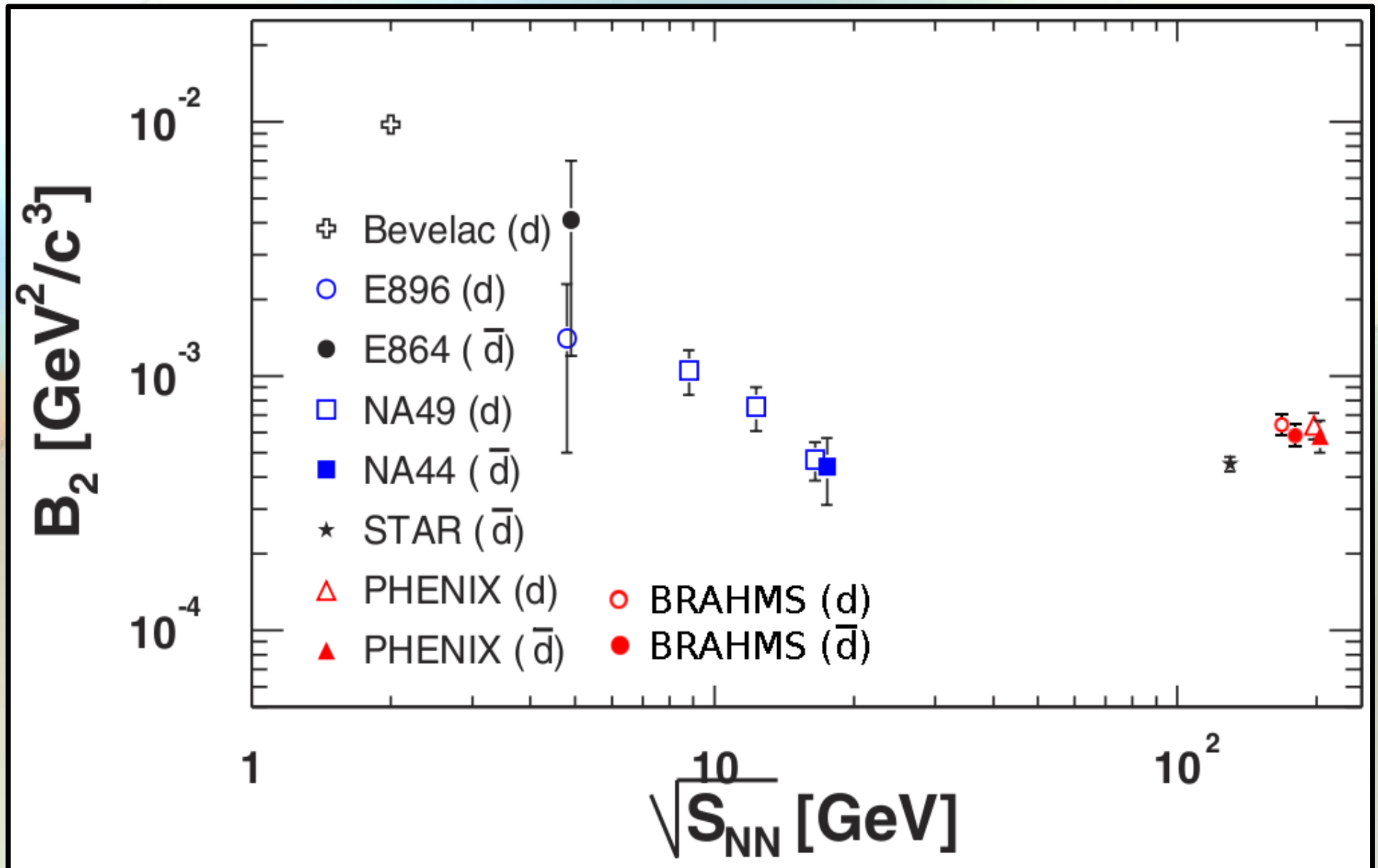


B_2 comparison to PHENIX



BRAHMS PRELIMINARY

Energy dependency



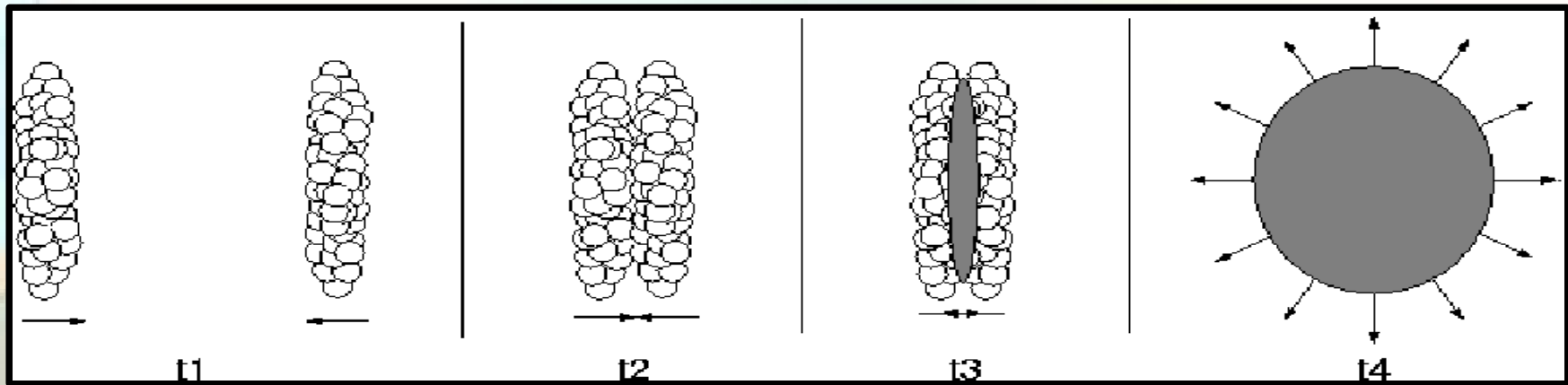
Coalescence Summary

- B_2 increases as a function of p_T at $y \sim 0$ and $y \sim 1$.
- B_2 is constant within errors in the rapidity range $y \sim [0; 3]$, indicating that source sizes are comparable at these rapidities.
- The decrease of B_2 as a function of collision energy is not observed at post RHIC energies.
- These results are due to being submitted for publication early 2009

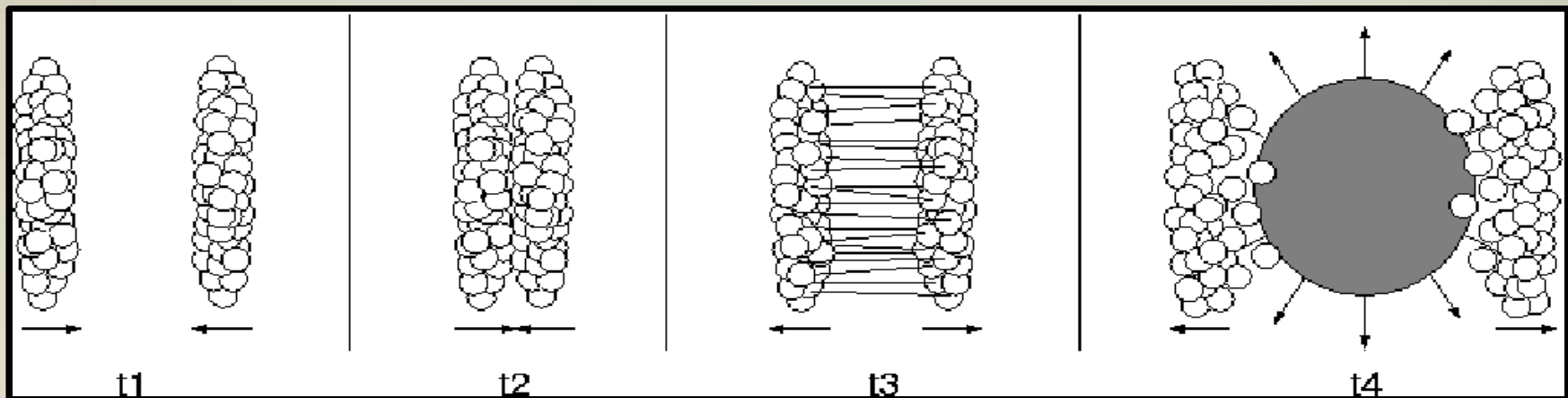
Nuclear stopping I

- Collision scenarios:

- Landau: Full stopping. Many baryons at midrapidity.



- Bjorken: Transparency. No baryons at midrapidity



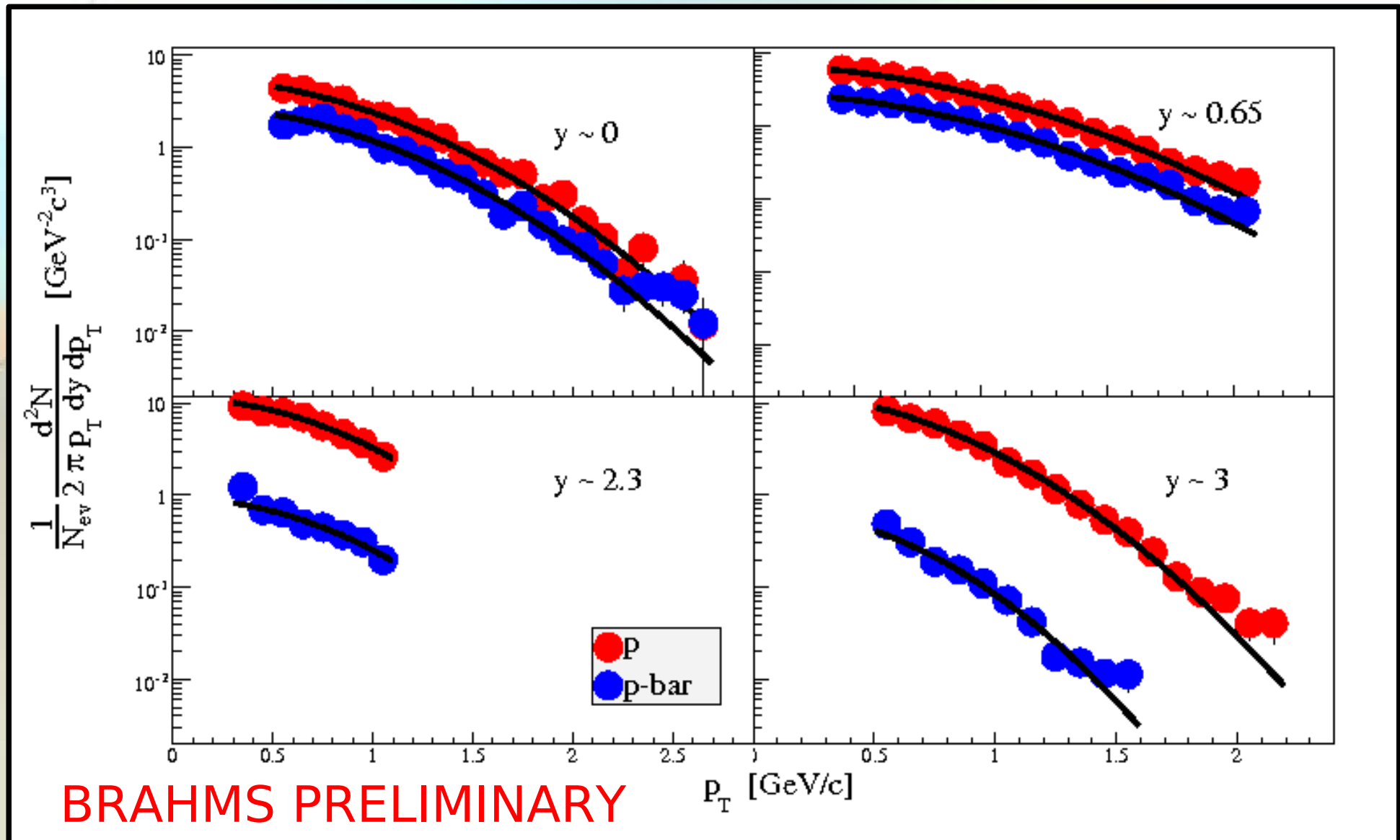
Nuclear Stopping II

- Quantify stopping by the rapidity loss:

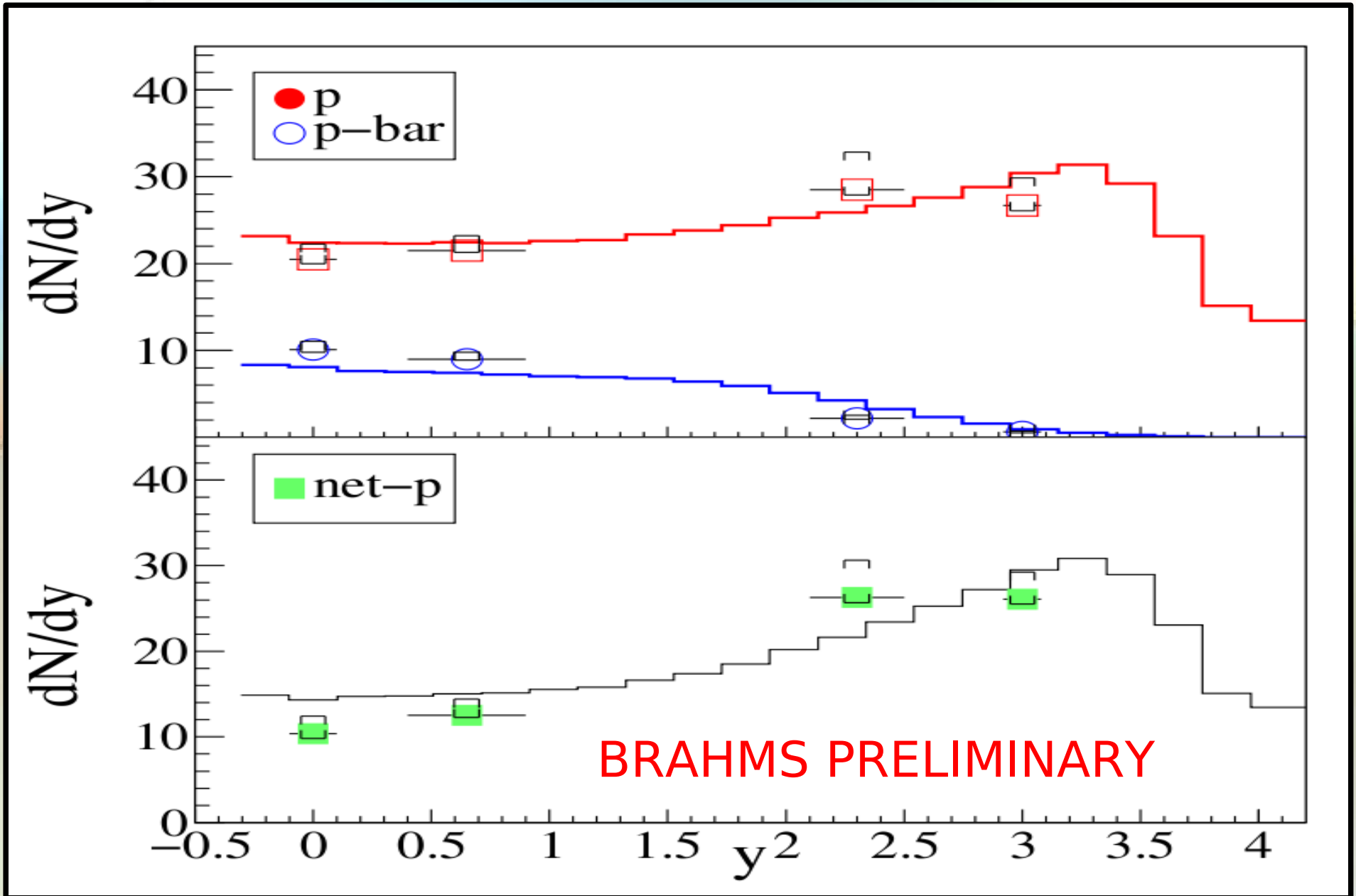
$$\delta y = y_{beam} - \langle y \rangle = y_{beam} - \frac{2}{N_{part}} \int_0^{y_{beam}} y \frac{dN_{net-baryons}}{dy} dy$$

- BRAHMS measures only charged hadrons, hence a conversion to baryons must be done.
- Baryon conservation is an important constraint.

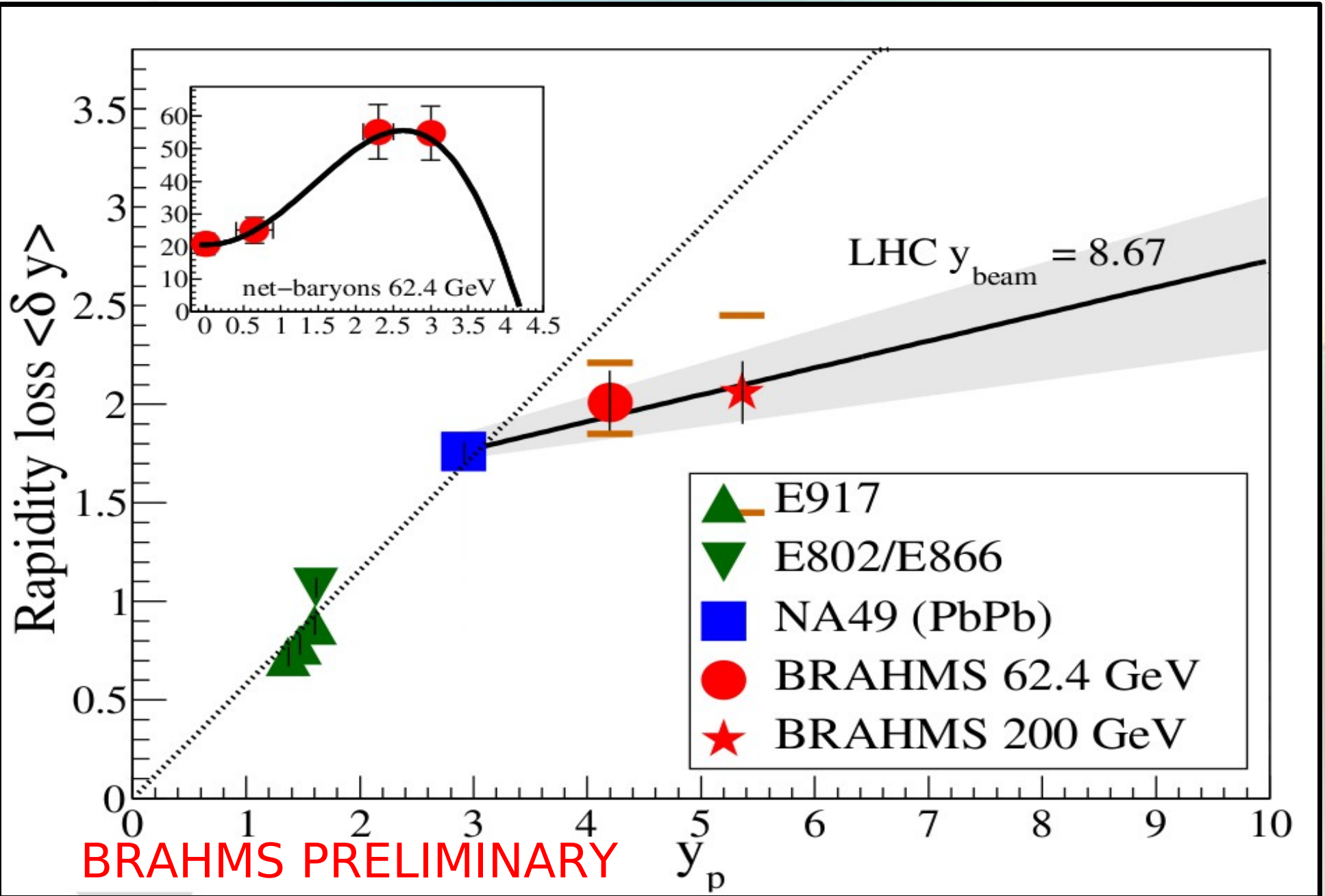
Spectra (0-10% central Au-Au@62.4 GeV)



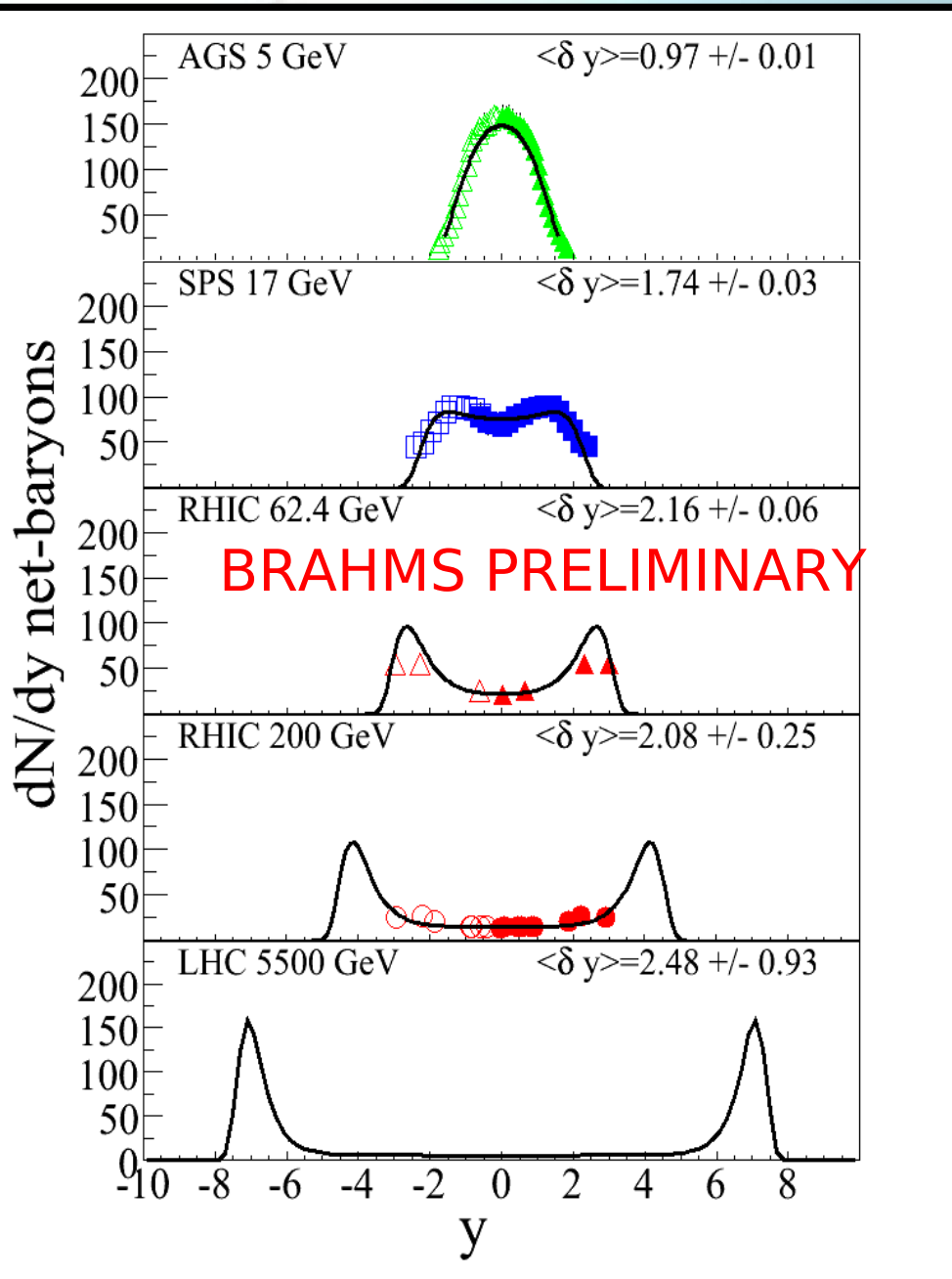
Yields - dN/dy



Rapidity loss



Net-baryons



- Fit: Bjorken inspired double gaussian in $p_z = m_T \sinh(y)$
- Baryon conversion factor:
 - $N_{\text{net-B}} \sim 2.5 N_{\text{net-p}}$ at AGS, SPS
 - $N_{\text{net-B}} \sim 2.1 N_{\text{net-p}}$ at RHIC, LHC
- Extrapolation to LHC done using simple straight line fits to μ, σ .

Nuclear Stopping Summary

- Stopping systematics might be used to predict LHC results or at least set limits.
- The linear scaling of rapidity loss is broken already at 62.4 GeV.
- This analysis is being submitted for publication before christmas 2008.

BRAHMS Collaboration

**I. C. Arsene¹², I. G. Bearden⁷, D. Beavis¹, S. Bekele¹², C. Besliu¹⁰, B. Budick⁶,
H. Bøgild⁷, C. Chasman¹, C. H. Christensen⁷, P. Christiansen⁷, H.H.Dalsgaard⁷, R. Debbe¹,
J. J. Gaardhøje⁷, K. Hagel⁸, H. Ito¹, A. Jipa¹⁰, E.B.Johnson¹¹, J. I. Jørdre⁹,
C. E. Jørgensen⁷, R. Karabowicz⁵, N. Katrynska⁵, E. J. Kim¹¹, T. M. Larsen⁷, J. H. Lee¹,
Y. K. Lee⁴, S. Lindahl¹², G. Løvholden¹², Z. Majka⁵, M. J. Murray¹¹, J. Natowitz⁸, C. Nygaard⁷
B. S. Nielsen⁸, D. Ouerdane⁸, D. Pal¹², F. Rami³, C. Ristea⁸, O. Ristea¹¹,
D. Röhrich⁹, B. H. Samset¹², S. J. Sanders¹¹, R. A. Scheetz¹, P. Staszal⁵,
T. S. Tveter¹², F. Videbæk¹, R. Wada⁸, H. Yang⁹, Z. Yin⁹, I. S. Zgura²**

1. Brookhaven National Laboratory, Upton, New York, USA

2. Institute of Space Science, Bucharest - Magurele, Romania

3. Institut Pluridisciplinaire Hubert Curien et Université Louis Pasteur, Strasbourg, France

4. Johns Hopkins University, Baltimore, USA

5. M. Smoluchowski Institute of Physics, Jagiellonian University, Krakow, Poland

6. New York University, New York, USA

7. Niels Bohr Institute, University of Copenhagen, Copenhagen, Denmark

8. Texas A&M University, College Station, Texas, USA

9. University of Bergen, Department of Physics and Technology, Bergen, Norway

10. University of Bucharest, Romania

11. University of Kansas, Lawrence, Kansas, USA

12. University of Oslo, Department of Physics, Oslo, Norway